

MEMORY EFFECT ANALYSIS USING PIECEWISE CUBIC B-SPLINE OF TIME FRACTIONAL DIFFUSION EQUATION

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Abstract

The purpose of this work is to study the memory effect analysis of Caputo–Fabrizio time fractional diffusion equation by means of cubic B-spline functions. The Caputo–Fabrizio interpretation of fractional derivative involves a non-singular kernel that permits to describe some class of material heterogeneities and the effect of memory more effectively. The proposed numerical technique relies on finite difference approach and cubic B-spline functions for discretization along temporal and spatial grids, respectively. To ensure that the error does not amplify during computational process, stability analysis is performed. The described algorithm is second-order convergent along time and space directions. The computational competence of the scheme is tested through some numerical examples. The results reveal that the current scheme is reasonably efficient and reliable to be used for solving the subject problem.

Keywords: Diffusion Equation; Spline Interpolation; Caputo–Fabrizio Fractional Derivative; Cubic B-Spline Functions; Stability; Finite Difference Formulation; Convergence.

1. INTRODUCTION

The advent of fractional calculus is as old as that of its classical counterpart. However, in the last two decades, it has gained a remarkable research attraction as it can be applied for solving many physical and engineering problems.^{1–3} The fractional order differential equations have the potential to describe the dynamic processes more efficiently as compared to classical ones. Numerous applications are found in biomedical engineering, electrochemistry, hydrology, finance, rheology, probability theory, optics, signal processing and other disciplines of science and technology.^{4–12} Several interpretations for fractional order derivatives have been proposed and each one of them has its own advantages and disadvantages, the interested readers are referred to Refs. 13–18. In this paper, we explore the approximate solutions of the following time fractional diffusion equation (TFDE):

$${}_0^{CF}D_t^\gamma w(v, t) = \frac{\partial^2 w(v, t)}{\partial v^2} + q(v, t), \quad \gamma \in (0, 1),$$

$$v \in [a, b], \quad t \in [t_0, T], \quad (1)$$

with initial condition

$$w(v, t_0) = \psi(v) \quad (2)$$

and the end conditions

$$w(a, t) = \phi_1(t), \quad w(b, t) = \phi_2(t), \quad (3)$$

where ${}_0^{CF}D_t^\gamma w(v, t)$ represents the Caputo–Fabrizio fractional derivative (CFFD), ϕ_1 , ϕ_2 , ψ , q are assumed to be continuous functions and γ denotes the fractional order of time derivative. The development of improved and reliable techniques for solving linear and nonlinear fractional diffusion models has always remained a hot area of research. The diffusion equation has been studied by various authors using the Caputo time fractional derivative instead of first-order time derivative. Zhuang and Liu¹⁹ employed an implicit difference approximation for solving TFDE. An explicit difference-based algorithm has been presented by Murillo and Yuste²⁰ for solutions of the fractional diffusion and diffusion wave equations with Caputo fractional derivative (CFD). Sweilam *et al.*²¹ proposed the Crank–Nicolson scheme for numerical solution of TFDE. Sun *et al.*²² employed the semi systematic finite element scheme for solving a collection of TFDEs. Mustapha *et al.*²³ developed a discontinuous Petrov–Galerkin technique for approximate solutions of TFDEs. In Ref. 24, Esmaeili and Garrappa constructed a pseudo-spectral method for numerical outcomes of TFDE. Tuan *et al.*²⁵ studied the inverse problem for one-dimensional TFDE by means of a modified regularization scheme using the frequency domain.

Atangana and Alkahtani²⁶ presented the model using fractional derivative with non-singular kernel



of resistance, inductance and capacitance electrical circuit. Atangana and Araz²⁷ have presented new concept in calculus of piecewise derivative with classical and global derivatives as well as with singular and non-singular kernels using numerical algorithm based on the Newton polynomial for epidemiological, fractional orders and fractal dimensions problems. In Ref. 28, Atangana and Alqahtani developed a numerical approach for advection dispersion equation with time-space CFFD and studied its application to groundwater pollution equation. A differential equation involving CFFD and its application to mass-spring-damper system have been investigated by Al-Salti *et al.*²⁹ Goufo³⁰ established the model of Korteweg–de Vries–Burgers equation with CFFD. Mirza and Vieru³¹ used the Fourier and Laplace transforms to obtain fundamental solutions for two-dimensional advection–diffusion equation in Caputo–Fabrizio sense. Liu *et al.*³² developed and analyzed the quasi-linear time fractional parabolic equation with non-singular kernel using finite difference scheme. Kumar *et al.*³³ proposed an iterative method for numerical solution of modified Kawahara equation involving fractional operator with exponential kernel. Shaikh *et al.*³⁴ investigated the approximate solutions of Fitzhugh–Nagumo and Fisher equation involving CFFD using an iterative Laplace transform method.

In recent years, numerous numerical techniques based on B-spline interpolation have been presented by different authors for solving the fractional differential equations. In particular, the spline functions are found to be powerful tools in curve approximation due to their rich geometrical features. The authors in Refs. 35–37 employed cubic trigonometric B-spline technique for approximate solutions of time fractional sub-diffusion, diffusion-wave and Burgers' equations. Khalid *et al.*³⁸ presented a technique based on modified extended cubic B-spline (CBS) functions for Caputo time fractional diffusion-wave equation with damping and reaction terms. In Ref. 39, Mohyud-Din *et al.* assembled extended cubic B-splines to develop a difference scheme for approximate solutions of time fractional advection diffusion equation. Numerical outcomes of time fractional telegraph equation have been presented by Akram *et al.*⁴⁰ using an extended form of CBS functions. Khalid *et al.*⁴¹ proposed the application of non-polynomial fifth degree spline functions for computational study of fourth-order fractional ordinary differential equations containing product terms. The authors in Refs. 42 and 43 propounded

a computational scheme based on non-polynomial quintic spline for solving fourth-order time fractional sub-diffusion and super-diffusion equations.

In this paper, a computational method based on CBS functions has been proposed for approximate solutions of TFDE. The θ -weighted formulation is used to carry out the presented algorithm. This method relies on Caputo–Fabrizio derivative with finite difference scheme for temporal discretization and CBS functions for interpolation along spatial grid. The stability and convergence of current numerical approach have also been presented. To the best of the author's knowledge, the proposed algorithm for TFDE is novel and it has not been noted in literature yet.

This paper is based on following framework: Some definitions are given in Sec. 2. In Sec. 3, derivation of the numerical method based on CBS functions is presented. The stability of the algorithm is analyzed in Sec. 4. Section 5 describes the convergence analysis. To corroborate the efficiency and accuracy of the scheme, results of computational experiments are recorded in Sec. 6 and finally, the conclusion is presented in Sec. 7.

2. PRELIMINARIES

Definition 1 (Ref. 27). Let $u(t)$ and $f(t)$ be two continuous functions such that $f(t)$ is increasing and non-constant and $u(t)$ is differentiable, then a piecewise derivative in local way within the interval $[0, T]$ is

$${}^0PGD_t u(t) = \begin{cases} \frac{du(t)}{dt} & \text{if } t \in [0, t_0], \\ D_f u(t) & \text{if } t \in [t_0, T], \end{cases} \quad (4)$$

where $\frac{du(t)}{dt}$ and $D_f u(t)$ represent the classical derivative on $0 \leq t \leq t_0$ and global derivative on $t_0 \leq t \leq T$, respectively.

Definition 2 (Ref. 27). The piecewise derivative with classical and exponential decay kernel is defined within the interval $[0, T]$ as

$${}^0PCFD_t^\gamma u(t) = \begin{cases} u'(t) & \text{if } t \in [0, t_0], \\ {}^0CFD_t^\gamma u(t) & \text{if } t \in [t_0, T] \end{cases} \quad (5)$$

and

$${}^0PCFD_t^\gamma u(t) = \begin{cases} u'(t) & \text{if } t \in [0, t_0], \\ {}^0CFRD_t^\gamma u(t) & \text{if } t \in [t_0, T], \end{cases} \quad (6)$$

where ${}^0PCFD_t^\gamma u(t)$ represents classical derivative on $0 \leq t \leq t_0$ and CFFD on $t_0 \leq t \leq T$.

Definition 3 (Ref. 44). The CFFD $\frac{\partial^\gamma}{\partial t^\gamma} w(v, t)$ with non-singular kernel of order $\gamma \in (0, 1)$ is defined as

$$\begin{aligned} {}_0^{CF} D_t^\gamma w(v, t) &= \frac{\partial^\gamma}{\partial t^\gamma} w(v, t) \\ &= \frac{R(\gamma)}{1-\gamma} \int_0^t \frac{\partial}{\partial \kappa} w(v, \kappa) \\ &\quad \times \exp \left[-\frac{\gamma}{1-\gamma} (t-\kappa) \right] d\kappa, \\ &\quad \gamma \in (0, 1), \end{aligned} \tag{7}$$

where the normalization function $R(\gamma)$ satisfying the condition $R(0) = R(1) = 1$.

Definition 4 (Ref. 45). Parseval's identity is defined as

$$\sum_{m=-\infty}^{\infty} |\hat{y}(m)|^2 = \int_a^b |\tilde{y}(v)|^2 dv, \quad \tilde{y} \in L^2[a, b], \tag{8}$$

where $\hat{y}(m) = \int_a^b \tilde{y}(v) e^{2\pi i m v} dv$ is the Fourier transform for every integer m .

3. THE DERIVATION OF NUMERICAL METHOD

Let the time domain $[0, T]$ be divided into M equal subintervals of the length $\Delta t = \frac{T}{M}$ using the knots $0 = t_0 < t_1 < \dots < t_M = T$, where $t_n = n\Delta t$, $n = 0, 1, \dots, M$ and the spatial domain $[a, b]$ be divided into N subintervals of equal length $h = \frac{b-a}{N}$ using the knots $a = v_0 < v_1 < \dots < v_N = b$, where $v_r = v_0 + rh$, $r = 0, 1, \dots, N$. For $w(v, t)$, the CBS approximation $W(v, t)$ can be assumed as

$$W(v, t) = \sum_{r=-1}^{N+1} \alpha_r(t) S_r(v), \tag{9}$$

where $\alpha_r(t)$, time-dependent control points, are to be computed and $S_r(v)$, the CBS functions, are described as⁴⁶

$$S_r(v) = \frac{1}{6h^3} \begin{cases} (v - v_{r-2})^3 & \text{if } v \in [v_{r-2}, v_{r-1}), \\ \varsigma & \text{if } v \in [v_{r-1}, v_r), \\ \chi & \text{if } v \in [v_r, v_{r+1}), \\ (v_{r+2} - v)^3 & \text{if } v \in [v_{r+1}, v_{r+2}), \\ 0 & \text{otherwise,} \end{cases} \tag{10}$$

where

$$\varsigma = h^3 + 3h^2(v - v_{r-1}) + 3h(v - v_{r-1})^2$$

$$\begin{aligned} &-3(v - v_{r-1})^3 \\ \chi &= h^3 + 3h^2(v_{r+1} - v) + 3h(v_{r+1} - v)^2 \\ &-3(v_{r+1} - v)^3. \end{aligned}$$

The CBS functions have a lot of geometrical properties such as symmetry, non-negativity, convex hull property, geometric invariability, local support and the partition of unity.⁴⁶ Moreover, $S_{-1}, S_0, \dots, S_{N+1}$ have been composed in such a way that they can serve as basis for space interval $[a, b]$. The relations (9) and (10) provide the following approximations:

$$\begin{cases} (W)_r = \left(\frac{1}{6}\right) \alpha_{r-1} + \left(\frac{4}{6}\right) \alpha_r + \left(\frac{1}{6}\right) \alpha_{r+1}, \\ (W_v)_r = \left(\frac{1}{2h}\right) \alpha_{r+1} + \left(-\frac{1}{2h}\right) \alpha_{r-1}, \\ (W_{vv})_r = \left(\frac{1}{h^2}\right) \alpha_{r-1} + \left(-\frac{2}{h^2}\right) \alpha_r \\ \quad + \left(\frac{1}{h^2}\right) \alpha_{r+1}. \end{cases} \tag{11}$$

Substituting (9) into (1), we get

$$\frac{\partial^\gamma W(v, t)}{\partial t^\gamma} = \frac{\partial^2 W(v, t)}{\partial v^2} + q(v, t). \tag{12}$$

The Caputo–Fabrizio fractional time derivative involved in (12) is discretized at $t = t_{n+1}$ as

$$\begin{aligned} \frac{\partial^\gamma}{\partial t^\gamma} W(v, t_{n+1}) &= \frac{1}{1-\gamma} \int_0^{t_{n+1}} \frac{\partial}{\partial \kappa} W(v, \kappa) \\ &\quad \times \exp \left[-\frac{\gamma}{1-\gamma} (t_{n+1} - \kappa) \right] d\kappa, \\ &= \frac{1}{1-\gamma} \sum_{z=0}^n \int_{t_z}^{t_{z+1}} \frac{\partial}{\partial \kappa} W(v, \kappa) \\ &\quad \times \exp \left[-\frac{\gamma}{1-\gamma} (t_{n+1} - \kappa) \right] d\kappa. \end{aligned} \tag{13}$$

Through forward difference formulation, equation (13) is modified as

$$\begin{aligned} \frac{\partial^\gamma}{\partial t^\gamma} W(v, t_{n+1}) &= \frac{1}{1-\gamma} \sum_{z=0}^n \frac{W(v, t_{z+1}) - W(v, t_z)}{\Delta t} \\ &\quad \times \int_{t_z}^{t_{z+1}} \exp \left[-\frac{\gamma}{1-\gamma} (t_{n+1} - \kappa) \right] d\kappa + \sigma_{\Delta t}^{n+1} \end{aligned}$$

$$B = \begin{pmatrix} 0 & 0 & 0 \\ p_1 & p_2 & p_1 \\ & p_1 & p_2 & p_1 \\ & & \ddots & \ddots & \ddots \\ & & & p_1 & p_2 & p_1 \\ & & & & p_1 & p_2 & p_1 \\ & & & & & 0 & 0 & 0 \end{pmatrix},$$

$$\alpha^n = \begin{pmatrix} \alpha_{-1}^n \\ \alpha_0^n \\ \alpha_1^n \\ \vdots \\ \alpha_{N-1}^n \\ \alpha_N^n \\ \alpha_{N+1}^n \end{pmatrix}$$

and

$$Q = \begin{pmatrix} \phi_1^{n+1} \\ \beta q_0^{n+1} \\ \beta q_1^{n+1} \\ \vdots \\ \beta q_{N-1}^{n+1} \\ \beta q_N^{n+1} \\ \phi_2^{n+1} \end{pmatrix}.$$

Before using (19), the initial vector $\alpha^0 = [\alpha_{-1}^0, \alpha_0^0, \dots, \alpha_{N+1}^0]^T$ is acquired by utilizing the initial conditions as

$$\begin{cases} (W_v)_r^0 = \psi'(v_r), & r = 0, \\ (W_r)_r^0 = \psi(v_r), & r = 0, 1, \dots, N, \\ (W_v)_r^0 = \psi'(v_r), & r = N. \end{cases} \quad (20)$$

The above system of equations is represented in matrix form as

$$C\alpha^0 = D, \quad (21)$$

where,

$$C = \begin{pmatrix} -p_3 & 0 & p_3 \\ p_1 & p_2 & p_1 \\ & p_1 & p_2 & p_1 \\ & & \ddots & \ddots & \ddots \\ & & & p_1 & p_2 & p_1 \\ & & & & p_1 & p_2 & p_1 \\ & & & & & -p_3 & 0 & p_3 \end{pmatrix}$$

and

$$D = \begin{pmatrix} \psi'(v_0) \\ \psi(v_0) \\ \psi(v_1) \\ \vdots \\ \psi(v_{N-1}) \\ \psi(v_N) \\ \psi'(v_N) \end{pmatrix}.$$

Equation (21) is easily solvable for α^0 using a suitable numerical algorithm. Mathematica 12 is used to carry out all numerical computations.

4. THE STABILITY ANALYSIS

A numerical scheme is accepted to be stable when the error does not increase during the computational procedure.⁴⁸ Here, stability of the scheme is analysed using the Fourier method.⁴⁶ Assume that ρ_r^n and $\tilde{\rho}_r^n$ represent the growth factor and its estimation, respectively. The error ζ_r^n can be represented as

$$\zeta_r^n = \rho_r^n - \tilde{\rho}_r^n, \quad r = 1(1)N - 1, \quad n = 0(1)M.$$

To keep the simplicity, we describe the stability of the proposed approach for the linear force-free form of (18) as

$$\begin{aligned} & (p_1 - \beta p_4)\zeta_{r-1}^{n+1} + (p_2 - \beta p_5)\zeta_r^{n+1} + (p_1 - \beta p_4)\zeta_{r+1}^{n+1} \\ & = (p_1\zeta_{r-1}^n + p_2\zeta_r^n + p_1\zeta_{r+1}^n) \\ & - \sum_{z=1}^n l_z [p_1(\zeta_{r-1}^{n-z+1} - \zeta_{r-1}^{n-z}) + p_2(\zeta_r^{n-z+1} \\ & - \zeta_r^{n-z}) + p_1(\zeta_{r+1}^{n-z+1} - \zeta_{r+1}^{n-z})]. \end{aligned} \quad (22)$$

From initial and end conditions, we can write

$$\zeta_r^0 = \psi(v_r), \quad r = 1, 2, \dots, N \quad (23)$$

and

$$\zeta_0^n = \phi_1(t_n), \quad \zeta_N^n = \phi_2(t_n), \quad n = 0, 1, \dots, M. \quad (24)$$

We introduce grid function

$$\zeta^n = \begin{cases} \zeta_r^n, & v_r - \frac{h}{2} < v \leq v_r \\ +\frac{h}{2}, & r = 1, 2, \dots, N - 1, \\ 0, & a \leq v \leq a + \frac{h}{2} \text{ or } b - \frac{h}{2} \leq v \leq b. \end{cases} \quad (25)$$

The Fourier expansion for $\zeta^n(v)$ can be expressed as

$$\zeta^n(v) = \sum_{m=-\infty}^{\infty} \delta^n(m) e^{\frac{2\pi i m v}{b-a}}, \quad (26)$$

where

$$\delta^n(m) = \frac{1}{b-a} \int_a^b \zeta^n(v) e^{-\frac{2\pi i m v}{b-a}} dv, \quad n = 0, 1, 2, \dots, M \quad (27)$$

and

$$\zeta^n = [\zeta_1^n, \zeta_2^n, \dots, \zeta_{N-1}^n]^T.$$

Applying $\|\cdot\|_2$ norm, we achieve

$$\begin{aligned} \|\zeta^n\|_2 &= \sqrt{h \sum_{r=1}^{N-1} |\zeta_r^n|^2} \\ &= \left(\int_a^{a+\frac{h}{2}} |\zeta^n|^2 dv + \sum_{r=1}^{N-1} \int_{v_r-\frac{h}{2}}^{v_r+\frac{h}{2}} |\zeta^n|^2 dv \right. \\ &\quad \left. + \int_{b-\frac{h}{2}}^b |\zeta^n|^2 dv \right)^{\frac{1}{2}} \\ &= \left(\int_a^b |\zeta^n|^2 dv \right)^{\frac{1}{2}}. \end{aligned}$$

Through Parseval's identity (8), we get⁴⁵

$$\int_a^b |\zeta^n|^2 dv = \sum_{m=-\infty}^{\infty} |\delta^n(m)|^2.$$

Hence, we acquire

$$\|\zeta^n\|_2^2 = \sum_{m=-\infty}^{\infty} |\delta^n(m)|^2. \quad (28)$$

Presume that (22)–(24) possess a solution in the Fourier form as

$$\zeta_r^n = \delta^n e^{i\epsilon r h}, \quad (29)$$

where ϵ is a real number and $i = \sqrt{-1}$. Putting (29) in (22), we achieve

$$\begin{aligned} &(p_1 - \beta p_4) \delta^{n+1} e^{-i\epsilon h} + (p_2 - \beta p_5) \delta^{n+1} \\ &\quad + (p_1 - \beta p_4) \delta^{n+1} e^{i\epsilon h} \\ &= (p_1 \delta^n e^{-i\epsilon h} + p_2 \delta^n + p_1 \delta^n e^{i\epsilon h}) \\ &\quad - \sum_{z=1}^n l_z [p_1 (\delta^{n-z+1} e^{-i\epsilon h} - \delta^{n-z} e^{-i\epsilon h}) \end{aligned}$$

$$\begin{aligned} &+ p_2 (\delta^{n-z+1} - \delta^{n-z}) \\ &\quad + p_1 (\delta^{n-z+1} e^{i\epsilon h} - \delta^{n-z} e^{i\epsilon h})]. \quad (30) \end{aligned}$$

Utilizing the relation $e^{i\epsilon h} + e^{-i\epsilon h} = 2\cos(\epsilon h)$, we acquire

$$\delta^{n+1} = \frac{1}{\eta} \left[\delta^n - \sum_{z=1}^n l_z (\delta^{n-z+1} - \delta^{n-z}) \right], \quad (31)$$

where $\eta = 1 + \frac{12\beta \sin^2(\epsilon h/2)}{h^2(3-2\sin^2(\epsilon h/2))}$. Obviously $\eta \geq 1$.

Lemma 5. Assume that δ^n is the solution for (31), then $|\delta^n| \leq |\delta^0|$, $n = 0, 1, \dots, T \times M$.

Proof. To examine this, mathematical induction is used. For $n = 0$, (31) implies

$$|\delta^1| = \frac{1}{\eta} |\delta^0| \leq |\delta^0|, \quad \eta \geq 1.$$

Assume that $|\delta^n| \leq |\delta^0|$ for $n = 1, 2, \dots, T \times (M-1)$, then

$$\begin{aligned} |\delta^{n+1}| &\leq \frac{1}{\eta} |\delta^n| - \frac{1}{\eta} \sum_{z=1}^n l_z (|\delta^{n-z+1}| - |\delta^{n-z}|) \\ &\leq \frac{1}{\eta} |\delta^0| - \frac{1}{\eta} \sum_{z=1}^n l_z (|\delta^0| - |\delta^0|) \\ &\leq |\delta^0|. \quad \square \end{aligned}$$

Theorem 6. The scheme (18) is unconditionally stable.

Proof. Employing the expression (28) and Lemma 5, we achieve

$$\|\zeta^n\|_2 \leq \|\zeta^0\|_2, \quad n = 0, 1, \dots, M.$$

Thus, the presented scheme is stable unconditionally. \square

5. THE CONVERGENCE ANALYSIS

To investigate convergence of the proposed method, we follow the procedure described in Ref. 49. Firstly, some useful theorems and lemmas are brought forward.^{50,51}

Theorem 7. Assuming $w(v, t)$, q belong to $C^4[a, b]$ and $C^2[a, b]$, respectively, and $\Psi = \{a = v_0, v_1, \dots, v_N = b\}$ is the partition of $[a, b]$ s.t. $v_r = a + rh, r = 0, 1, \dots, N$. If $\tilde{W}(v, t)$ is the CBS approximation

to (1) at points $v_r \in \Psi$, then there exists Ω_r not depending on h , we have

$$\|D^r(w(v, t) - \tilde{W}(v, t))\|_\infty \leq \Omega_r h^{4-r}, \quad \forall t \geq 0, \quad r = 0, 1, 2. \quad (32)$$

Lemma 8. The CBS set $\{S_{-1}, S_0, \dots, S_{N+1}\}$ presented in (10) satisfies the inequality

$$\sum_{r=-1}^{N+1} |S_r(v)| \leq \frac{5}{3}, \quad 0 \leq v \leq 1. \quad (33)$$

Proof. We can write the following relation using the triangular inequality:

$$\left| \sum_{r=-1}^{N+1} S_r(v) \right| \leq \sum_{r=-1}^{N+1} |S_r(v)|.$$

At any knot v_r , we acquire

$$\begin{aligned} \sum_{r=-1}^{N+1} |S_r(v)| &= |S_{r-1}(v_r)| + |S_r(v_r)| + |S_{r+1}(v_r)| \\ &= \frac{1}{6} + \frac{4}{6} + \frac{1}{6} = 1 < \frac{5}{3}. \end{aligned}$$

Moreover, for $v \in [v_r, v_{r+1}]$, we get

$$\begin{aligned} |S_r(v)| &\leq \frac{4}{6}, \quad |S_{r-1}(v)| \leq \frac{1}{6}, \\ |S_{r+1}(v)| &\leq \frac{4}{6} \quad \text{and} \quad |S_{r+2}(v)| \leq \frac{1}{6}. \end{aligned}$$

Hence, for any point $v_r \leq v \leq v_{r+1}$, we attain

$$\begin{aligned} \sum_{r=-1}^{N+1} |S_r(v)| &= |S_{r-1}(v)| + |S_r(v)| + |S_{r+1}(v)| \\ &\quad + |S_{r+2}(v)| \leq \frac{5}{3}. \quad \square \end{aligned}$$

Theorem 9. The numerical approximation $W(v, t)$ to the exact solution $w(v, t)$ for TFDE exists. Furthermore, if q is a member of $C^2[0, 1]$, then

$$\|w(v, t) - W(v, t)\|_\infty \leq \tilde{\Omega} h^2, \quad \forall t \geq 0, \quad (34)$$

where $\tilde{\Omega} > 0$ is free of h and h is appropriately small.

Proof. We suppose that $\tilde{W}(v, t) = \sum_{r=-1}^{N+1} a_r^n(t) S_r(v)$ be the approximated CBS for $W(v, t)$. By implementing the triangular inequality, we obtain

$$\begin{aligned} \|w(v, t) - W(v, t)\|_\infty &\leq \|w(v, t) - \tilde{W}(v, t)\|_\infty \\ &\quad + \|\tilde{W}(v, t) - W(v, t)\|_\infty. \end{aligned}$$

Making use of Theorem 7, we get

$$\begin{aligned} \|w(v, t) - W(v, t)\|_\infty &\leq \Omega_0 h^4 + \|\tilde{W}(v, t) - W(v, t)\|_\infty. \quad (35) \end{aligned}$$

$Lw(v_r, t) = LW(v_r, t) = q(v_r, t)$, $r = 0(1)N$ are the collocation conditions. Assume that

$$L\tilde{W}(v_r, t) = \tilde{q}(v_r, t), \quad r = 0(1)N.$$

Thus, for any time stage n , the difference $L(\tilde{W}(v_r, t) - W(v_r, t))$ can be stated as

$$\begin{aligned} &(p_1 - \beta p_4) \chi_{r-1}^{n+1} + (p_2 - \beta p_5) \chi_r^{n+1} \\ &\quad + (p_1 - \beta p_4) \chi_{r+1}^{n+1} \\ &= (p_1 \chi_{r-1}^n + p_2 \chi_r^n + p_1 \chi_{r+1}^n) + \frac{\beta}{h^2} \xi_r^{n+1} \\ &\quad - \sum_{z=1}^n l_z [p_1 (\chi_{r-1}^{n-z+1} - \chi_{r-1}^{n-z}) \\ &\quad + p_2 (\chi_r^{n-z+1} - \chi_r^{n-z}) \\ &\quad + p_1 (\chi_{r+1}^{n-z+1} - \chi_{r+1}^{n-z})]. \quad (36) \end{aligned}$$

The end conditions can be described as

$$p_1 \chi_{r-1}^{n+1} + p_2 \chi_r^{n+1} + p_1 \chi_{r+1}^{n+1} = 0, \quad r = 0, N,$$

where

$$\chi_r^n = \alpha_r^n - a_r^n, \quad r = -1 : 0 : N + 1$$

and

$$\xi_r^n = h^2 [q_r^n - \tilde{q}_r^n], \quad r = 0 : 1 : N.$$

From inequality (32), we achieve

$$|\xi_r^n| = h^2 |q_r^n - \tilde{q}_r^n| \leq \Omega h^4.$$

Define $\xi^n = \max\{|\xi_r^n|; 0 \leq r \leq N\}$, $E_r^n = |\chi_r^n|$ and $E^n = \max\{E_r^n; 0 \leq r \leq N\}$.

When $n = 0$, Eq. (36) becomes

$$\begin{aligned} &(p_1 - \beta p_4) \chi_{r-1}^1 + (p_2 - \beta p_5) \chi_r^1 + (p_1 - \beta p_4) \chi_{r+1}^1 \\ &= (p_1 \chi_{r-1}^0 + p_2 \chi_r^0 + p_1 \chi_{r+1}^0) + \frac{\beta}{h^2} \xi_r^1, \end{aligned}$$

where $r = 0 : 1 : N$. Through initial condition, $E^0 = 0$:

$$(p_2 - \beta p_5) \chi_r^1 = -(p_1 - \beta p_4) (\chi_{r-1}^1 + \chi_{r+1}^1) + \frac{\beta}{h^2} \xi_r^1.$$

Taking norms of χ_r^1 , ξ_r^1 and adequately small grid spacing h , we have

$$E_r^1 \leq \frac{3\beta}{h^2 + 12\beta} \Omega h^4, \quad r = 0 : 1 : N.$$

The values of E_{-1}^1 and E_{N+1}^1 are obtained from the end conditions:

$$E_{-1}^1 \leq \frac{15\beta}{h^2 + 12\beta} \Omega h^4,$$

$$E_{N+1}^1 \leq \frac{15\beta}{h^2 + 12\beta} \Omega h^4,$$

which implies

$$E^1 \leq \Omega_1 h^2, \tag{37}$$

where Ω_1 does not rely on h .

Now, mathematical induction on n is used. Assume that $E_r^s \leq \Omega_s h^2$ is true for $s = 1, 2, \dots, n$ and $\Omega = \max\{\Omega_s : s = 0, 1, \dots, n\}$, then from Eq. (36), we acquire

$$\begin{aligned} & (p_1 - \beta p_4) \chi_{r-1}^{n+1} + (p_2 - \beta p_5) \chi_r^{n+1} + (p_1 - \beta p_4) \\ & \quad \times \chi_{r+1}^{n+1} \\ &= [(l_0 - l_1)(p_1 \chi_{r-1}^n + p_2 \chi_r^n + p_1 \chi_{r+1}^n) \\ & \quad + (l_1 - l_2)(p_1 \chi_{r-1}^{n-1} + p_2 \chi_r^{n-1} + p_1 \chi_{r+1}^{n-1}) \\ & \quad + \dots + (l_{n-1} - l_n)(p_1 \chi_{r-1}^1 + p_2 \chi_r^1 \\ & \quad + p_1 \chi_{r+1}^1) + l_n(p_1 \chi_{r-1}^0 + p_2 \chi_r^0 + p_1 \chi_{r+1}^0)] \\ & \quad + \frac{\beta}{h^2} \xi_r^{n+1}. \end{aligned}$$

Again, applying the norms on χ_r^{n+1} and ξ_r^{n+1} , we get

$$E_r^{n+1} \leq \frac{3\Omega h^4}{h^2 + 12\beta} \left(\beta + \sum_{z=0}^{n-1} (l_z - l_{z+1}) \right).$$

Similarly, we obtain the values of E_{-1}^{n+1} and E_{N+1}^{n+1} from the end conditions:

$$E_{-1}^{n+1} \leq \frac{15\Omega h^4}{h^2 + 12\beta} \left(\beta + \sum_{z=0}^{n-1} (l_z - l_{z+1}) \right),$$

$$E_{N+1}^{n+1} \leq \frac{15\Omega h^4}{h^2 + 12\beta} \left(\beta + \sum_{z=0}^{n-1} (l_z - l_{z+1}) \right).$$

Hence, for all n , we acquire

$$E^{n+1} \leq \Omega h^2. \tag{38}$$

In particular,

$$\tilde{W}(v, t) - W(v, t) = \sum_{r=-1}^{N+1} (a_r(t) - \alpha_r(t)) S_r(v).$$

Therefore, from inequality (38) and Lemma 8, we get

$$\|\tilde{W}(v, t) - W(v, t)\|_\infty \leq \frac{5}{3} \Omega h^2. \tag{39}$$

Employing (39), the inequality (35) becomes

$$\|w(v, t) - W(v, t)\|_\infty \leq \Omega_0 h^4 + \frac{5}{3} \Omega h^2 = \tilde{\Omega} h^2,$$

where $\tilde{\Omega} = \Omega_0 h^2 + \frac{5}{3} \Omega$. □

Theorem 10. *The TFDE is convergent with initial and end conditions.*

Proof. Assume that $w(v, t)$ and $W(v, t)$ are analytical and numerical solutions for the TFDE, respectively. As a result of relation (15) and the preceding theorem, there exist arbitrary constants $\tilde{\Omega}$ and ϑ such that

$$\|w(v, t) - W(v, t)\|_\infty \leq \tilde{\Omega} h^2 + \vartheta (\Delta t)^2.$$

Consequently, in spatial and temporal directions, the presented method is second-order convergent. □

6. NUMERICAL EXAMPLES AND DISCUSSION

In this section, numerical outcomes of some experiments using presented method are reported. The computational efficiency of the technique is checked through the error norms $e_2(h, \Delta t)$ and $e_\infty(h, \Delta t)$ as

$$\begin{aligned} e_2(h, \Delta t) &= \|w(v_r, t) - W(v_r, t)\|_2 \\ &= \sqrt{h \sum_{r=0}^N |w(v_r, t) - W(v_r, t)|^2} \end{aligned}$$

and

$$\begin{aligned} e_\infty(h, \Delta t) &= \|w(v_r, t) - W(v_r, t)\|_\infty \\ &= \max_{0 \leq r \leq N} |w(v_r, t) - W(v_r, t)|. \end{aligned}$$

Moreover, the convergence order in spatial and temporal directions is computed as⁴⁷

$$\log_2 \left| \frac{e_\infty(h, \Delta t)}{e_\infty(\frac{h}{2}, \Delta t)} \right| \text{ and } \log_2 \left| \frac{e_\infty(h, \Delta t)}{e_\infty(h, \frac{\Delta t}{2})} \right|.$$

Example 11. Consider the TFDE⁴⁷

$$\frac{\partial^\gamma w(v, t)}{\partial t^\gamma} = \frac{\partial^2 w(v, t)}{\partial v^2} + q(v, t), \quad v \in [0, 1], \quad t \in [0, 1],$$

with initial and end conditions

$$w(v, 0) = e^{\gamma v}$$

and

$$w(0, t) = e^t, \quad w(1, t) = e^{\gamma+t},$$

where $q(v, t) = -e^{\gamma v+t} (e^{-\frac{1}{1-\gamma}t} + \gamma^2 - 1)$.

Table 1 Absolute Error When $t = 1$, $\gamma = 0.1$, $\Delta t = 0.01$ and $N = 10$ of Example 11.

v	Exact Solution	Approximate Solution	Absolute Error
0.1	2.7456010150169163	2.7456010773020605	6.22851×10^{-8}
0.2	2.7731947639642978	2.7731948744068227	1.10443×10^{-7}
0.3	2.8010658346990790	2.8010659795360200	1.44837×10^{-7}
0.4	2.8292170143515600	2.8292171800365207	1.65685×10^{-7}
0.5	2.8576511180631640	2.8576512911203724	1.73057×10^{-7}
0.6	2.8863709892679585	2.8863711561470584	1.66879×10^{-7}
0.7	2.9153794999769970	2.9153796469070320	1.46930×10^{-7}
0.8	2.9446795510655240	2.9446796639065513	1.12841×10^{-7}
0.9	2.9742740725630656	2.9742741366538230	6.40908×10^{-8}

Table 2 Absolute Error When $t = 1$, $N = 20$, $\gamma = 0.3$ and $\Delta t = 0.01$ of Example 11.

v	Exact Solution	Approximate Solution	Absolute Error
0.1	2.8010658346990790	2.8010659598562370	1.25157×10^{-7}
0.2	2.8863709892679585	2.8863712124252214	2.23157×10^{-7}
0.3	2.9742740725630656	2.9742743670020357	2.94439×10^{-7}
0.4	3.0648542032930024	3.0648545423483250	3.39055×10^{-7}
0.5	3.1581929096897670	3.1581932663632020	3.56673×10^{-7}
0.6	3.2543742028896707	3.2543745494587630	3.46569×10^{-7}
0.7	3.3534846525490236	3.3534849601651650	3.07616×10^{-7}
0.8	3.4556134647626755	3.4556137030331793	2.38271×10^{-7}
0.9	3.5608525623555205	3.5608526989041350	1.36549×10^{-7}

Table 3 Error Norms for Different γ Choices with $\Delta t = 0.01$ and $N = 32$ for Example 11.

t	$e_\infty(h, \Delta t)$		$e_2(h, \Delta t)$	
	$\gamma = 0.3$	$\gamma = 0.5$	$\gamma = 0.3$	$\gamma = 0.5$
0.2	3.50778×10^{-8}	3.98217×10^{-7}	2.57441×10^{-8}	2.91972×10^{-7}
0.4	1.69987×10^{-7}	1.20488×10^{-7}	1.24707×10^{-7}	8.79828×10^{-8}
0.6	3.25632×10^{-7}	1.62955×10^{-7}	2.38850×10^{-7}	1.19994×10^{-7}
0.8	5.07259×10^{-7}	4.61873×10^{-7}	3.72020×10^{-7}	3.39143×10^{-7}
1.0	7.21229×10^{-7}	7.86892×10^{-7}	5.28880×10^{-7}	5.77290×10^{-7}

The exact solution is $w(v, t) = e^{\gamma v+t}$. Absolute errors are reported in Tables 1 and 2 for Example 11 when $t = 1$, $\Delta t = 0.01$, $\gamma = 0.1, 0.3$ and $N = 10, 20$, respectively. At different time intervals, error norms are listed in Table 3. Table 4 shows the comparison of maximum error with the results given in Ref. 47 at $t = 1$ of Example 11 for different γ choices. Table 5 compares convergence order and error by setting $\gamma = 0.5$ at $t = 1$. The plots of exact and computational results at various time stages for $N = 16, 100$, $\gamma = 0.5$ and $\Delta t = 0.01$ are depicted in Fig. 1. Figure 2 shows the accuracy of

the current scheme through 3D graphs of analytical and numerical solutions. At $t = 1$, 2D and 3D error profiles are displayed in Fig. 3. Figure 4 illustrates the numerical and exact solutions when $N = 20, 100$ and $\Delta t = 0.01$ by setting different choices of γ at $t = 1$. In comparison to the scheme presented in Ref. 47, the proposed scheme is found to be more efficient and accurate. The numerical solutions using CBS approximations when $\gamma = 0.5, 0.75$, $t = 0.5, 1$, $\Delta t = 0.01$, $0 \leq v \leq 1$ and $h = 0.02, 0.05$ are

Table 4 Maximum Error's Comparison When $t = 1$ for Example 11.

h	Δt	FDM ⁴⁷			Proposed Method		
		$\gamma = 0.25$	$\gamma = 0.75$	$\gamma = 0.95$	$\gamma = 0.25$	$\gamma = 0.75$	$\gamma = 0.95$
0.05	0.05	1.78606×10^{-5}	3.00280×10^{-4}	2.05777×10^{-3}	1.73054×10^{-5}	2.41731×10^{-4}	1.88707×10^{-3}
	0.02	3.54460×10^{-6}	3.30745×10^{-5}	4.06247×10^{-4}	2.53244×10^{-6}	1.37860×10^{-5}	2.34313×10^{-4}
	0.01	9.84836×10^{-7}	4.05129×10^{-5}	1.66279×10^{-4}	4.21817×10^{-7}	1.88081×10^{-5}	5.83444×10^{-6}
0.02	0.05	1.76317×10^{-5}	2.76236×10^{-4}	1.98642×10^{-3}	1.75429×10^{-5}	2.66772×10^{-4}	1.95910×10^{-3}
	0.02	2.85917×10^{-6}	4.82200×10^{-5}	3.34114×10^{-4}	2.76922×10^{-6}	3.87260×10^{-5}	3.06600×10^{-4}
	0.01	7.48608×10^{-7}	1.56190×10^{-5}	9.40322×10^{-5}	6.58504×10^{-7}	6.11960×10^{-6}	6.64896×10^{-5}
0.01	0.05	1.76032×10^{-5}	2.72688×10^{-4}	1.97617×10^{-3}	1.75809×10^{-5}	2.70322×10^{-4}	1.96935×10^{-3}
	0.02	2.82620×10^{-6}	4.46604×10^{-5}	3.23797×10^{-4}	2.80371×10^{-6}	4.22867×10^{-5}	3.16918×10^{-4}
	0.01	7.15013×10^{-7}	1.20570×10^{-5}	8.37043×10^{-5}	6.92481×10^{-7}	9.68207×10^{-6}	7.68187×10^{-5}

given by

$$W(v, 0.5) = \begin{cases} 1.64872 + v(0.824363 + (0.206086 + 0.0345207v)v) & \text{if } v \in [0.00, 0.02) \\ 1.64872 + v(0.824364 + (0.206065 + 0.0348676v)v) & \text{if } v \in [0.02, 0.04) \\ 1.64872 + v(0.824365 + (0.206023 + 0.035218v)v) & \text{if } v \in [0.04, 0.06) \\ 1.64872 + v(0.824369 + (0.205959 + 0.0355719v)v) & \text{if } v \in [0.06, 0.08) \\ 1.64872 + v(0.824376 + (0.205873 + 0.0359294v)v) & \text{if } v \in [0.08, 0.10) \\ 1.64872 + v(0.824387 + (0.205765 + 0.0362904v)v) & \text{if } v \in [0.10, 0.12) \\ 1.64872 + v(0.824403 + (0.205634 + 0.0366551v)v) & \text{if } v \in [0.12, 0.14) \\ 1.64872 + v(0.824424 + (0.205479 + 0.0370235v)v) & \text{if } v \in [0.14, 0.16) \\ \vdots & \\ 1.64847 + v(0.826491 + (0.199424 + 0.0434469v)v) & \text{if } v \in [0.46, 0.48) \\ 1.64842 + v(0.826793 + (0.198795 + 0.0438835v)v) & \text{if } v \in [0.48, 0.50) \\ 1.64837 + v(0.827123 + (0.198133 + 0.0443245v)v) & \text{if } v \in [0.50, 0.52) \\ 1.6483 + v(0.827485 + (0.197438 + 0.0447699v)v) & \text{if } v \in [0.52, 0.54) \\ \vdots & \\ 1.64557 + v(0.838914 + (0.181261 + 0.0525374v)v) & \text{if } v \in [0.84, 0.86) \\ 1.64523 + v(0.840085 + (0.179899 + 0.0530654v)v) & \text{if } v \in [0.86, 0.88) \\ 1.64487 + v(0.841324 + (0.178491 + 0.0535987v)v) & \text{if } v \in [0.88, 0.90) \\ 1.64447 + v(0.842633 + (0.177036 + 0.0541373v)v) & \text{if } v \in [0.90, 0.92) \\ 1.64405 + v(0.844015 + (0.175535 + 0.0546814v)v) & \text{if } v \in [0.92, 0.94) \\ 1.64359 + v(0.845471 + (0.173985 + 0.0552309v)v) & \text{if } v \in [0.94, 0.96) \\ 1.6431 + v(0.847006 + (0.172387 + 0.0557859v)v) & \text{if } v \in [0.96, 0.98) \\ 1.64258 + v(0.848621 + (0.170738 + 0.0563466v)v) & \text{if } v \in [0.98, 1.00) \end{cases} \tag{40}$$

Table 5 Error Norm's Comparison for Example 11 When $t = 1$ and $\gamma = 0.5$.

h	Δt	FDM ⁴⁷		Proposed Method		
		$e_\infty(h, \Delta t)$	Order	$e_\infty(h, \Delta t)$	$e_2(h, \Delta t)$	Order
0.05	0.02	1.61978×10^{-5}	...	6.01029×10^{-6}	4.41463×10^{-6}	...
	0.01	7.87199×10^{-6}	1.04100	2.32171×10^{-6}	1.70262×10^{-6}	1.37225
0.02	0.02	1.19452×10^{-5}	...	1.03118×10^{-5}	7.55533×10^{-6}	...
	0.01	3.59933×10^{-6}	1.73063	1.96501×10^{-6}	1.43997×10^{-6}	2.39169
0.01	0.02	1.13327×10^{-5}	...	1.09244×10^{-5}	8.00390×10^{-6}	...
	0.01	2.98650×10^{-6}	1.92396	2.57792×10^{-6}	1.88881×10^{-6}	2.08327

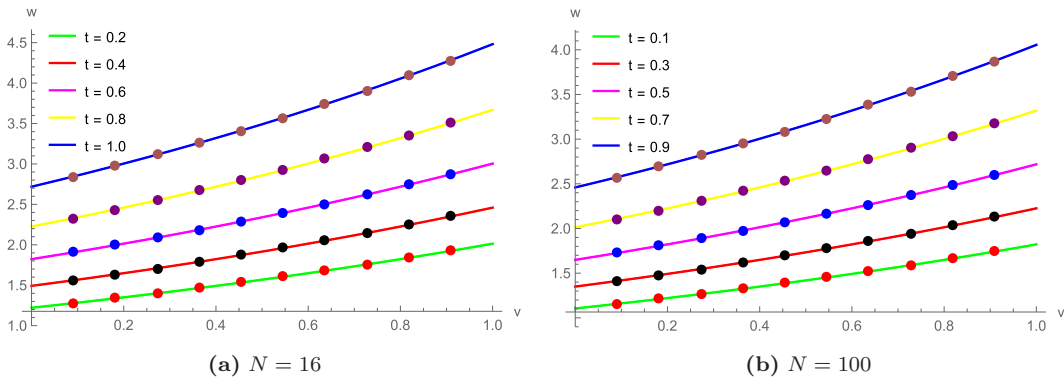


Fig. 1 Exact and approximate solutions at various time levels with $\Delta t = 0.01$ and $\gamma = 0.5$ for Example 11.

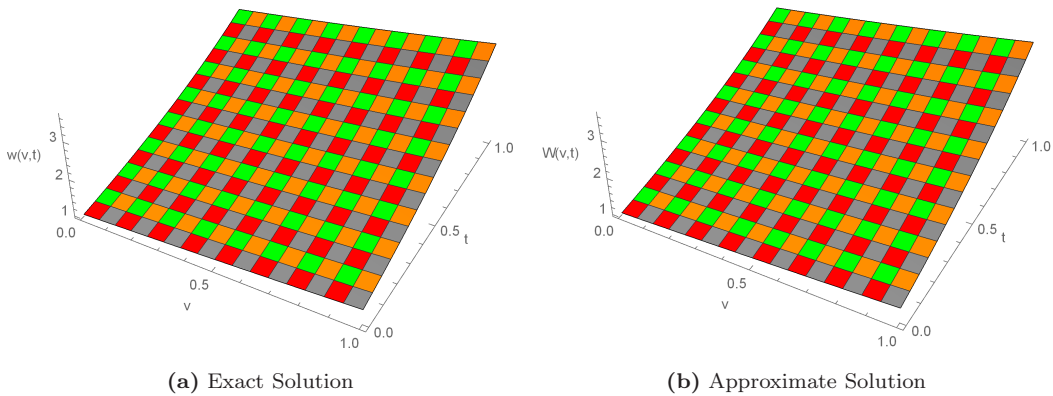


Fig. 2 3D exact and approximate solutions with $N = 100$, $\gamma = 0.3$, $t = 1$, $v \in [0, 1]$ and $\Delta t = 0.01$ for Example 11.

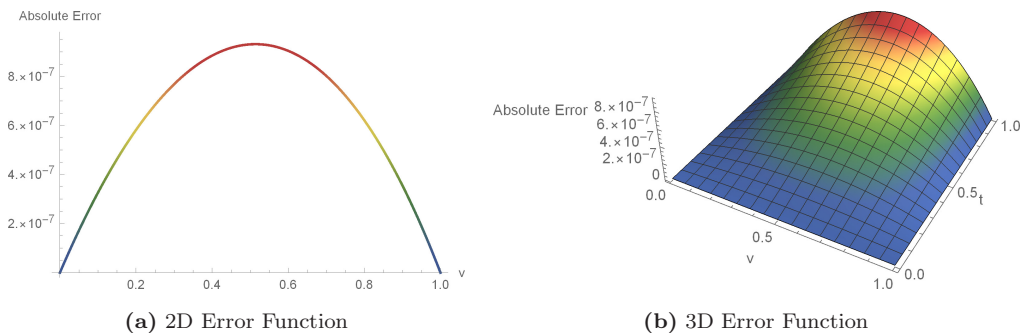


Fig. 3 2D and 3D error profiles with $N = 100$, $\gamma = 0.3$, $v \in [0, 1]$, $t = 1$ and $\Delta t = 0.01$ for Example 11.

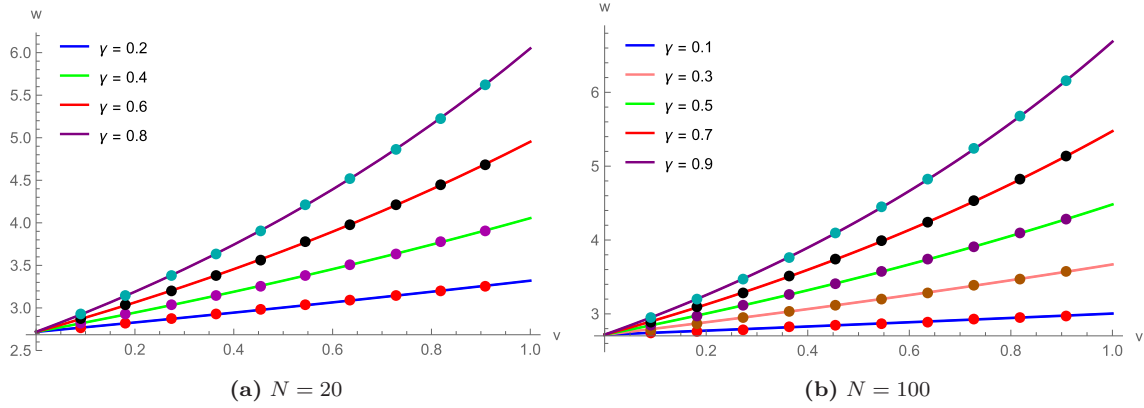


Fig. 4 Exact and approximate solutions at various γ values for Example 11 with $\Delta t = 0.01$ and $t = 1$.

and

$$W(v, 1) = \begin{cases} 2.71828 + v(2.03864 + (0.76448 + 0.19474v)v) & \text{if } v \in [0.00, 0.05) \\ 2.71828 + v(2.0387 + (0.763367 + 0.202182v)v) & \text{if } v \in [0.05, 0.10) \\ 2.71827 + v(2.03893 + (0.761049 + 0.209909v)v) & \text{if } v \in [0.10, 0.15) \\ 2.71825 + v(2.03947 + (0.757439 + 0.217931v)v) & \text{if } v \in [0.15, 0.20) \\ 2.71818 + v(2.04047 + (0.752442 + 0.22626v)v) & \text{if } v \in [0.20, 0.25) \\ 2.71804 + v(2.04209 + (0.745956 + 0.234907v)v) & \text{if } v \in [0.25, 0.30) \\ 2.7178 + v(2.04452 + (0.737876 + 0.243885v)v) & \text{if } v \in [0.30, 0.35) \\ 2.7174 + v(2.04794 + (0.72809 + 0.253205v)v) & \text{if } v \in [0.35, 0.40) \\ 2.71678 + v(2.05259 + (0.716478 + 0.262882v)v) & \text{if } v \in [0.40, 0.45) \\ 2.71587 + v(2.05869 + (0.702915 + 0.272928v)v) & \text{if } v \in [0.45, 0.50) \\ 2.71456 + v(2.06651 + (0.687269 + 0.283359v)v) & \text{if } v \in [0.50, 0.55) \\ 2.71276 + v(2.07634 + (0.669401 + 0.294188v)v) & \text{if } v \in [0.55, 0.60) \\ 2.71033 + v(2.08848 + (0.649164 + 0.305431v)v) & \text{if } v \in [0.60, 0.65) \\ 2.70713 + v(2.10328 + (0.626403 + 0.317103v)v) & \text{if } v \in [0.65, 0.70) \\ 2.70297 + v(2.12109 + (0.600954 + 0.329221v)v) & \text{if } v \in [0.70, 0.75) \\ 2.69766 + v(2.14232 + (0.572646 + 0.341803v)v) & \text{if } v \in [0.75, 0.80) \\ 2.69098 + v(2.1674 + (0.541297 + 0.354865v)v) & \text{if } v \in [0.80, 0.85) \\ 2.68265 + v(2.1968 + (0.506715 + 0.368427v)v) & \text{if } v \in [0.85, 0.90) \\ 2.67238 + v(2.23101 + (0.4687 + 0.382506v)v) & \text{if } v \in [0.90, 0.95) \\ 2.65985 + v(2.27059 + (0.427039 + 0.397124v)v) & \text{if } v \in [0.95, 1.00). \end{cases} \tag{41}$$

Example 12. Consider the TFDE⁴⁷

$$\frac{\partial^\gamma w(v, t)}{\partial t^\gamma} = \frac{\partial^2 w(v, t)}{\partial v^2} + q(v, t), \quad v \in [0, 1], \quad t \in [0, 5],$$

with initial and end conditions

$$w(v, 0) = \cos \gamma v$$

and

$$w(0, t) = 1 + \sin \frac{\gamma}{1-\gamma} t,$$

$$w(1, t) = \cos \gamma + \sin \frac{\gamma}{1-\gamma} t,$$

where $q(v, t) = \gamma^2 \cos \gamma v + \frac{1}{2-2\gamma} (\cos(\frac{\gamma}{1-\gamma} t) + \sin(\frac{\gamma}{1-\gamma} t) - \exp(-\frac{\gamma}{1-\gamma} t))$.

The exact solution is $w(v, t) = \cos \gamma v + \sin \frac{\gamma}{1-\gamma} t$. The numerical results and absolute errors for Example 12 at various spatial grid points with $N = 100$, $\gamma = 0.3, 0.5$ and $\Delta t = 0.01$ at $t = 1, 4$ are shown in Tables 6 and 7. Error norms for various γ values are displayed in Table 8 at different time stages.

Table 9 compares the maximum relative error at various time stages with $\gamma = 0.5$. The maximum absolute error $e_\infty(h, \Delta t)$ for $\gamma = 0.5$, $\Delta t = 0.01$ and $h = 0.02, 0.01$ is reported in Tables 10 and 11. Tables 12 and 13 compare error and convergence order with those obtained in Ref. 47 along temporal and spatial grids, respectively. The close agreement between exact solutions and numerical results of the proposed scheme at various time levels can be seen in Fig. 5. Comparison of computational and exact solutions is displayed in Fig. 6 for $N = 32$, $\Delta t = 0.01$, $\gamma = 0.5$, $t = 5$ and $v \in [0, 1]$. At $t = 5$, 2D and 3D error profiles are demonstrated in Fig. 7. Figure 8 exhibits the behavior of exact values and computational results for various γ choices. The approximate solutions using CBS approximations with $\gamma = 0.5, 0.3$, $0 \leq v \leq 1$, $t = 1, 5$, $h = 0.02, 0.05$ and $\Delta t = 0.01$ are given by

$$W(v, 1) = \begin{cases} 1.84147 + v(-3.45765 \times 10^{-6} + (-0.124997 + 0.000103466v)v) & \text{if } v \in [0.00, 0.02) \\ 1.84147 + v(-3.20762 \times 10^{-6} + (-0.12501 + 0.000311818v)v) & \text{if } v \in [0.02, 0.04) \\ 1.84147 + v(-2.20769 \times 10^{-6} + (-0.125035 + 0.000520139v)v) & \text{if } v \in [0.04, 0.06) \\ 1.84147 + v(4.16136 \times 10^{-8} + (-0.125072 + 0.000728407v)v) & \text{if } v \in [0.06, 0.08) \\ 1.84147 + v(4.03896 \times 10^{-6} + (-0.125122 + 0.000936602v)v) & \text{if } v \in [0.08, 0.10) \\ 1.84147 + v(0.000010282 + (-0.125185 + 0.0011447v)v) & \text{if } v \in [0.10, 0.12) \\ 1.84147 + v(0.000019267 + (-0.125259 + 0.00135269v)v) & \text{if } v \in [0.12, 0.14) \\ 1.84147 + v(0.0000314886 + (-0.125347 + 0.00156054v)v) & \text{if } v \in [0.14, 0.16) \\ \vdots & \\ 1.84135 + v(0.00105985 + (-0.1284 + 0.00485084v)v) & \text{if } v \in [0.46, 0.48) \\ 1.84132 + v(0.00119974 + (-0.128692 + 0.00505322v)v) & \text{if } v \in [0.48, 0.50) \\ 1.8413 + v(0.00135115 + (-0.128994 + 0.00525511v)v) & \text{if } v \in [0.50, 0.52) \\ 1.84127 + v(0.00151449 + (-0.129309 + 0.00545646v)v) & \text{if } v \in [0.52, 0.54) \\ \vdots & \\ 1.84019 + v(0.00605082 + (-0.135782 + 0.00859047v)v) & \text{if } v \in [0.84, 0.86) \\ 1.84007 + v(0.00647105 + (-0.13627 + 0.00877986v)v) & \text{if } v \in [0.86, 0.88) \\ 1.83994 + v(0.00690901 + (-0.136768 + 0.00896838v)v) & \text{if } v \in [0.88, 0.90) \\ 1.83981 + v(0.00736493 + (-0.137274 + 0.009156v)v) & \text{if } v \in [0.90, 0.92) \\ 1.83966 + v(0.00783901 + (-0.13779 + 0.00934271v)v) & \text{if } v \in [0.92, 0.94) \\ 1.83951 + v(0.00833146 + (-0.138314 + 0.00952848v)v) & \text{if } v \in [0.94, 0.96) \\ 1.83934 + v(0.00884245 + (-0.138846 + 0.0097133v)v) & \text{if } v \in [0.96, 0.98) \\ 1.83917 + v(0.00937215 + (-0.139386 + 0.00989715v)v) & \text{if } v \in [0.98, 1.00) \end{cases} \quad (42)$$

Table 6 Absolute Error at Different Spatial Grid Points When $t = 1$, $\Delta t = 0.01$, $\gamma = 0.3$ and $N = 100$ for Example 12.

v	Exact Solution	Approximate Solution	Absolute Error
0.1	1.4151218887420396	1.4151218789931579	9.74888×10^{-9}
0.2	1.4137723949282561	1.4137723777162572	1.72120×10^{-8}
0.3	1.4115245880050462	1.4115245655286313	2.24764×10^{-8}
0.4	1.4083804908469182	1.4083804652424539	2.56045×10^{-8}
0.5	1.4043429329290942	1.4043429062946715	2.66344×10^{-8}
0.6	1.3994155477811734	1.3994155222002440	2.55809×10^{-8}
0.7	1.3936027697172002	1.3936027472820223	2.24352×10^{-8}
0.8	1.3869098298450815	1.3869098126803014	1.71648×10^{-8}
0.9	1.3793427513589425	1.3793427416455640	9.71338×10^{-9}

and

$$W(v, 5) = \begin{cases} 1.84079 + v(-1.60668 \times 10^{-6} + (-0.0449992 + 0.0000336533v)v) & \text{if } v \in [0.00, 0.05) \\ 1.84079 + v(-1.10042 \times 10^{-6} + (-0.0450093 + 0.000101155v)v) & \text{if } v \in [0.05, 0.10) \\ 1.84079 + v(9.23936 \times 10^{-7} + (-0.0450295 + 0.000168633v)v) & \text{if } v \in [0.10, 0.15) \\ 1.84079 + v(5.47617 \times 10^{-6} + (-0.0450599 + 0.000236073v)v) & \text{if } v \in [0.15, 0.20) \\ 1.84079 + v(0.0000135627 + (-0.0451003 + 0.000303461v)v) & \text{if } v \in [0.20, 0.25) \\ 1.84079 + v(0.000026185 + (-0.0451508 + 0.00037078v)v) & \text{if } v \in [0.25, 0.30) \\ 1.84078 + v(0.0000443386 + (-0.0452113 + 0.000438016v)v) & \text{if } v \in [0.30, 0.35) \\ 1.84078 + v(0.0000690115 + (-0.0452818 + 0.000505153v)v) & \text{if } v \in [0.35, 0.40) \\ 1.84078 + v(0.000101183 + (-0.0453623 + 0.000572176v)v) & \text{if } v \in [0.40, 0.45) \\ 1.84077 + v(0.000141821 + (-0.0454526 + 0.000639071v)v) & \text{if } v \in [0.45, 0.50) \\ 1.84076 + v(0.000191885 + (-0.0455527 + 0.000705822v)v) & \text{if } v \in [0.50, 0.55) \\ 1.84075 + v(0.000252317 + (-0.0456626 + 0.000772414v)v) & \text{if } v \in [0.55, 0.60) \\ 1.84074 + v(0.000324049 + (-0.0457821 + 0.000838832v)v) & \text{if } v \in [0.60, 0.65) \\ 1.84074 + v(0.000324049 + (-0.0457821 + 0.000838832v)v) & \text{if } v \in [0.65, 0.70) \\ 1.8407 + v(0.000505053 + (-0.0460499 + 0.000971088v)v) & \text{if } v \in [0.70, 0.75) \\ 1.84067 + v(0.000616103 + (-0.046198 + 0.0010369v)v) & \text{if } v \in [0.75, 0.80) \\ 1.84063 + v(0.000742005 + (-0.0463554 + 0.00110247v)v) & \text{if } v \in [0.80, 0.85) \\ 1.84059 + v(0.000883599 + (-0.0465219 + 0.0011678v)v) & \text{if } v \in [0.85, 0.90) \\ 1.84055 + v(0.0010417 + (-0.0466976 + 0.00123286v)v) & \text{if } v \in [0.90, 0.95) \\ 1.84049 + v(0.00121711 + (-0.0468823 + 0.00129764v)v) & \text{if } v \in [0.95, 1.00). \end{cases} \tag{43}$$

Example 13. Consider the TFDE⁴⁷

$$\frac{\partial^\gamma w(v, t)}{\partial t^\gamma} = \frac{\partial^2 w(v, t)}{\partial v^2} + q(v, t), \quad v \in [0, 1], \quad t \in [0, 1],$$

Table 7 Absolute Error at Different Spatial Grid Points When $t = 4$, $\Delta t = 0.01$, $\gamma = 0.5$ and $N = 100$ for Example 12.

v	Exact Solution	Approximate Solution	Absolute Error
0.1	0.2419477650870381	0.2419477255887654	3.94983×10^{-8}
0.2	0.2382016699700976	0.2382015944086032	7.55615×10^{-8}
0.3	0.2319685826281140	0.2319684785434553	1.04085×10^{-7}
0.4	0.2232640825333134	0.2232639603306453	1.22203×10^{-7}
0.5	0.2121099264027165	0.2121097981611280	1.28242×10^{-7}
0.6	0.1985339938176778	0.1985338721281287	1.21690×10^{-7}
0.7	0.1825702175394507	0.1825701143524707	1.03187×10^{-7}
0.8	0.1642584986949569	0.1642584241588478	7.45361×10^{-8}
0.9	0.1436446070447487	0.1436445683149489	3.87298×10^{-8}

Table 8 Error Norms for Various γ Values When $N = 64$, $\Delta t = 0.01$ and $v \in [0, 1]$ for Example 12.

t	$e_\infty(h, \Delta t)$			$e_2(h, \Delta t)$		
	$\gamma = 0.1$	$\gamma = 0.2$	$\gamma = 0.3$	$\gamma = 0.1$	$\gamma = 0.2$	$\gamma = 0.3$
0.2	2.31281×10^{-10}	3.68650×10^{-9}	1.87269×10^{-8}	1.69456×10^{-10}	2.70135×10^{-9}	1.37254×10^{-8}
0.4	2.41132×10^{-10}	3.96647×10^{-9}	2.13709×10^{-8}	1.76668×10^{-10}	2.90608×10^{-9}	1.56598×10^{-8}
0.6	2.57047×10^{-10}	4.40948×10^{-9}	2.54827×10^{-8}	1.88314×10^{-10}	3.23030×10^{-9}	1.86701×10^{-8}
0.8	2.78889×10^{-10}	5.00732×10^{-9}	3.09262×10^{-8}	2.04304×10^{-10}	3.66797×10^{-9}	2.26564×10^{-8}
1.0	3.06537×10^{-10}	5.75181×10^{-9}	3.75631×10^{-8}	2.24549×10^{-10}	4.21307×10^{-9}	2.75167×10^{-8}

Table 9 Maximum Relative Error Comparison at Different Time Stages with $\gamma = 0.5$ for Example 12.

t	FDM ⁴⁷	Proposed Method
0.5	2.61714×10^{-8}	2.80767×10^{-7}
1.0	2.07059×10^{-7}	4.63313×10^{-7}
2.0	6.11440×10^{-7}	8.69958×10^{-7}
3.0	9.47515×10^{-7}	1.32111×10^{-6}
4.0	6.87512×10^{-6}	1.51711×10^{-6}

Table 10 $e_\infty(h, \Delta t)$ at Various Time Levels Where $h = 0.02$, $\Delta t = 0.01$ and $\gamma = 0.5$ for Example 12.

t	0.5	1.0	2.0	3.0	4.0
$e_\infty(h, \Delta t)$	4.06645×10^{-7}	8.38773×10^{-7}	1.63396×10^{-6}	1.46569×10^{-6}	3.15290×10^{-7}

Table 11 $e_\infty(h, \Delta t)$ at Various Time Levels Where $h = 0.01$, $\Delta t = 0.01$ and $\gamma = 0.5$ of Example 12.

t	0.5	1.0	2.0	3.0	4.0
$e_\infty(h, \Delta t)$	2.39962×10^{-7}	6.64741×10^{-7}	1.45188×10^{-6}	1.28013×10^{-6}	1.28242×10^{-7}

Table 12 Convergence Order for Example 12 in Temporal Direction Where $h = \frac{1}{100}$ and $\gamma = 0.5$.

Δt	FDM ⁴⁷		Proposed Method		
	$e_\infty(h, \Delta t)$	Order	$e_\infty(h, \Delta t)$	$e_2(h, \Delta t)$	Order
$\frac{1}{10}$	6.03850×10^{-5}	...	1.08453×10^{-4}	7.96216×10^{-5}	...
$\frac{1}{20}$	1.49327×10^{-5}	2.01572	2.70761×10^{-5}	1.98783×10^{-5}	2.00198
$\frac{1}{40}$	3.55985×10^{-6}	2.06859	6.72270×10^{-6}	4.93577×10^{-6}	2.00991
$\frac{1}{80}$	7.15987×10^{-7}	2.31381	1.63378×10^{-6}	1.19972×10^{-6}	2.04083

Table 13 Convergence Order for Example 12 in Spatial Direction Where $\Delta t = \frac{1}{100}$ and $\gamma = 0.5$.

h	FDM ⁴⁷		Proposed Method		
	$e_\infty(h, \Delta t)$	Order	$e_\infty(h, \Delta t)$	$e_2(h, \Delta t)$	Order
$\frac{1}{4}$	3.55335×10^{-5}	...	3.79958×10^{-5}	2.76733×10^{-5}	...
$\frac{1}{8}$	8.44998×10^{-6}	2.07216	8.68896×10^{-6}	6.33612×10^{-6}	2.12858
$\frac{1}{16}$	1.65880×10^{-6}	2.34881	1.35826×10^{-6}	9.86688×10^{-7}	2.67742
$\frac{1}{32}$	3.53690×10^{-7}	2.22958	4.74662×10^{-7}	3.51171×10^{-7}	1.51679

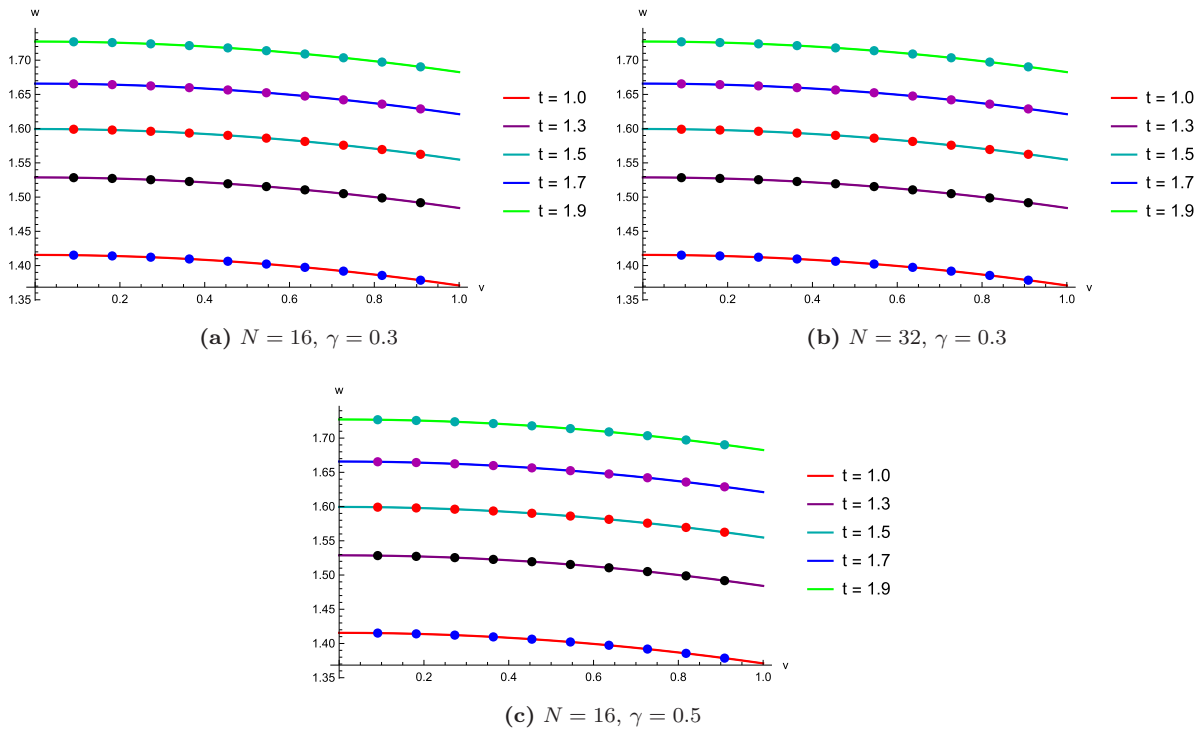


Fig. 5 Exact and approximate solutions at different time levels using $\Delta t = 0.01$ for Example 12.

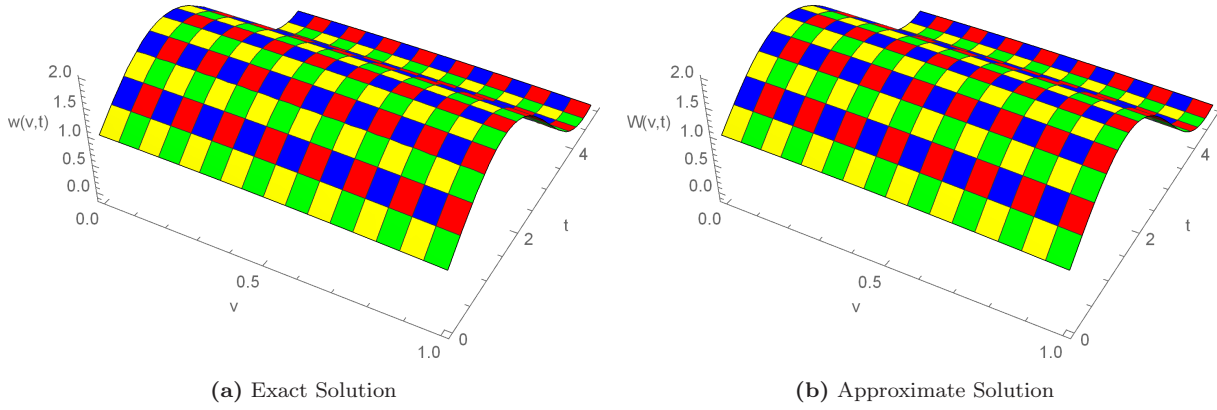


Fig. 6 3D exact and approximate solutions when $N = 32$, $\Delta t = 0.01$, $\gamma = 0.5$, $t = 5$ and $v \in [0, 1]$ for Example 12. and

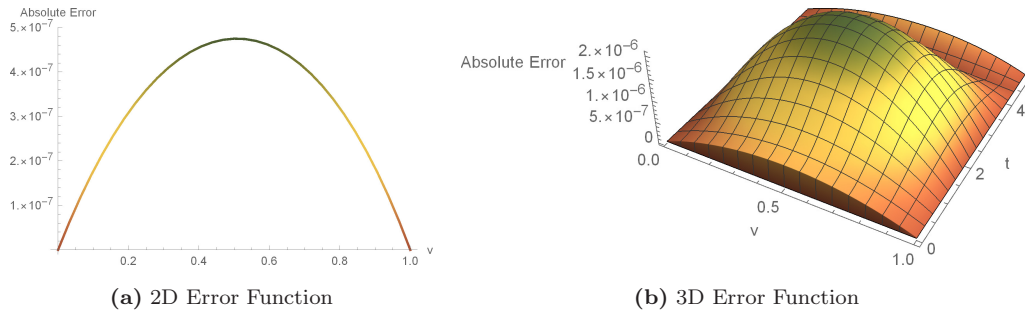


Fig. 7 2D and 3D error profiles when $N = 32$, $\Delta t = 0.01$, $\gamma = 0.5$, $t = 5$ and $v \in [0, 1]$ for Example 12.

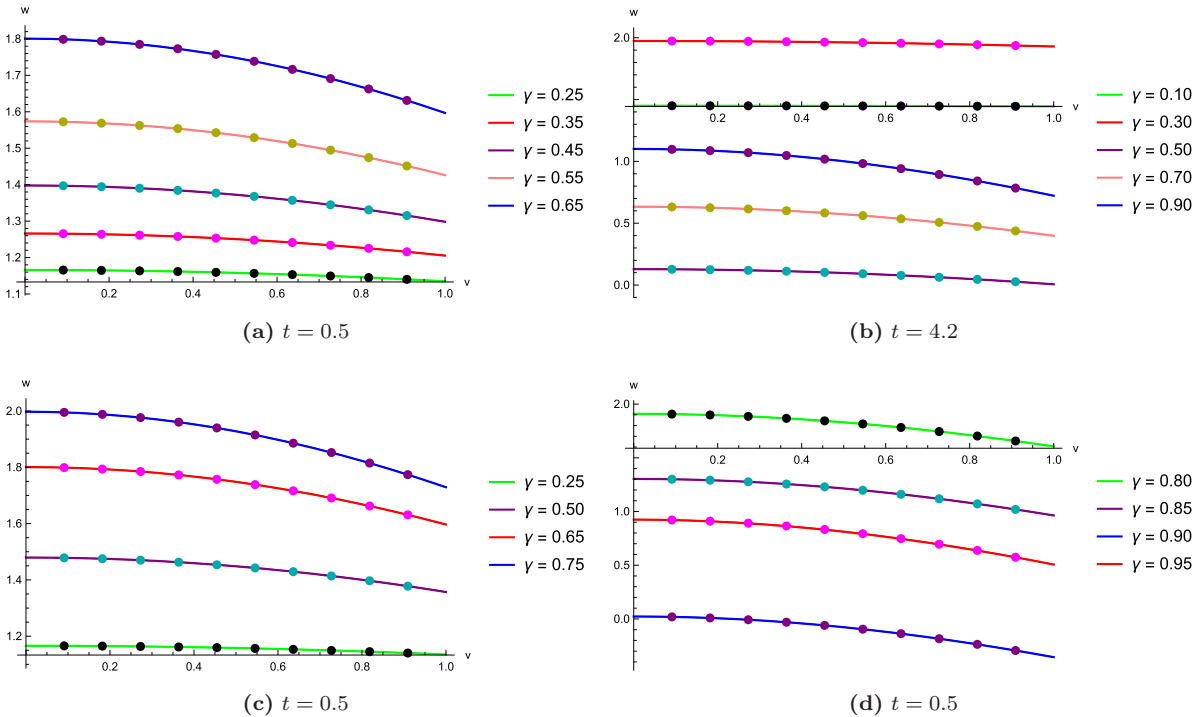


Fig. 8 Exact and approximate solutions at different γ values with $\Delta t = 0.01$ and $N = 16$ for Example 12.

with initial and end conditions

$$w(v, 0) = 1 + e^v$$

and

$$w(0, t) = 1 + e^{\frac{\gamma}{1-\gamma}t}, \quad w(1, t) = e + e^{\frac{\gamma}{1-\gamma}t},$$

where $q(v, t) = \frac{1}{1-\gamma} \sinh(\frac{\gamma}{1-\gamma}t) - e^v$.

The exact solution is $w(v, t) = e^v + e^{\frac{\gamma}{1-\gamma}t}$. The absolute errors for various values of v setting $\Delta t = 0.01$, $N = 80, 100$, $\gamma = 0.5, 0.3$ and $t = 1$ are reported in Tables 14 and 15 of Example 13. In Table 16, error norms are displayed for $N = 64$, $\Delta t = 0.01$, $\gamma = 0.6$ and $v \in [0, 1]$ at numerous time levels. Tables 17 and 18 expound the convergence

order comparison with those reported in Ref. 47 in temporal and spatial directions, respectively. Figure 9 displays the comparison of computational solution with exact results for Example 13 at various temporal grid points by setting $\gamma = 0.3, 0.5$ and $N = 32$. A space-time plot of computational and exact results is shown in Fig. 10 for $N = 90$, $\Delta t = 0.01$, $\gamma = 0.5$ and $t = 1$. At $t = 1$, 2D and 3D error profiles are exhibited in Fig. 11. Figure 12 displays approximate and exact solutions when $\Delta t = 0.01$, $N = 16$ and $t = 1$ corresponding to various choices of γ . The numerical solutions using CBS approximations when $\gamma = 0.6, 0.5$, $0 \leq v \leq 1$, $t = 0.5, 1$, $h = 0.05, 0.02$ and $\Delta t = 0.01$ are given by

$$W(v, 0.5) = \begin{cases} 3.117 + v(0.999883 + (0.499981 + 0.170882v)v) & \text{if } v \in [0.00, 0.05] \\ 3.117 + v(0.999948 + (0.498666 + 0.179646v)v) & \text{if } v \in [0.05, 0.10] \\ 3.11699 + v(1.00022 + (0.495902 + 0.188859v)v) & \text{if } v \in [0.10, 0.15] \\ 3.11696 + v(1.00088 + (0.491544 + 0.198545v)v) & \text{if } v \in [0.15, 0.20] \\ 3.11688 + v(1.0021 + (0.485434 + 0.208727v)v) & \text{if } v \in [0.20, 0.25] \\ 3.11671 + v(1.00411 + (0.477406 + 0.219432v)v) & \text{if } v \in [0.25, 0.30] \\ 3.1164 + v(1.00715 + (0.467278 + 0.230685v)v) & \text{if } v \in [0.30, 0.35] \\ 3.1159 + v(1.01149 + (0.454856 + 0.242515v)v) & \text{if } v \in [0.35, 0.40] \\ 3.1151 + v(1.01746 + (0.439932 + 0.254952v)v) & \text{if } v \in [0.40, 0.45] \\ 3.11391 + v(1.02541 + (0.422282 + 0.268026v)v) & \text{if } v \in [0.45, 0.50] \\ 3.11219 + v(1.03571 + (0.401664 + 0.281771v)v) & \text{if } v \in [0.50, 0.55] \\ 3.10979 + v(1.04883 + (0.377823 + 0.296221v)v) & \text{if } v \in [0.55, 0.60] \\ 3.10651 + v(1.06523 + (0.35048 + 0.311411v)v) & \text{if } v \in [0.60, 0.65] \\ 3.10212 + v(1.08547 + (0.319339 + 0.32738v)v) & \text{if } v \in [0.65, 0.70] \\ 3.09636 + v(1.11015 + (0.284084 + 0.344169v)v) & \text{if } v \in [0.70, 0.75] \\ 3.08892 + v(1.13994 + (0.244374 + 0.361818v)v) & \text{if } v \in [0.75, 0.80] \\ 3.07942 + v(1.17556 + (0.199845 + 0.380371v)v) & \text{if } v \in [0.80, 0.85] \\ 3.06744 + v(1.21784 + (0.150106 + 0.399877v)v) & \text{if } v \in [0.85, 0.90] \\ 3.05249 + v(1.26766 + (0.0947418 + 0.420382v)v) & \text{if } v \in [0.90, 0.95] \\ 3.03401 + v(1.32603 + (0.033305 + 0.441939v)v) & \text{if } v \in [0.95, 1.00] \end{cases} \quad (44)$$

Table 14 Absolute Error at Different Spatial Grid Points When $t = 1$, $N = 80$, $\gamma = 0.5$ and $\Delta t = 0.01$ of Example 13.

v	Exact Solution	Approximate Solution	Absolute Error
0.1	3.8234527465346930	3.8234527299243420	1.66104×10^{-8}
0.2	3.9396845866192150	3.9396845055345895	8.10846×10^{-8}
0.3	4.0681406360350480	4.0681404604812540	1.75554×10^{-7}
0.4	4.2101065261003150	4.2101062443559040	2.81744×10^{-7}
0.5	4.3670030991591740	4.3670027185348540	3.80624×10^{-7}
0.6	4.5404006288495540	4.5404001768060080	4.52044×10^{-7}
0.7	4.7320345359295220	4.7320340615631350	4.74366×10^{-7}
0.8	4.9438227569515130	4.9438223328607940	4.24091×10^{-7}
0.9	5.1778849396159945	5.1778846641651580	2.75451×10^{-7}

and

$$W(v, 1) = \begin{cases} 3.71828 + v(0.999987 + (0.49999 + 0.168344v)v) & \text{if } v \in [0.00, 0.02) \\ 3.71828 + v(0.999991 + (0.499786 + 0.171744v)v) & \text{if } v \in [0.02, 0.04) \\ 3.71828 + v(1.00001 + (0.49937 + 0.175214v)v) & \text{if } v \in [0.04, 0.06) \\ 3.71828 + v(1.00005 + (0.498733 + 0.178753v)v) & \text{if } v \in [0.06, 0.08) \\ 3.71828 + v(1.00011 + (0.497866 + 0.182364v)v) & \text{if } v \in [0.08, 0.10) \\ 3.71828 + v(1.00023 + (0.496761 + 0.186049v)v) & \text{if } v \in [0.10, 0.12) \\ 3.71827 + v(1.00039 + (0.495408 + 0.189807v)v) & \text{if } v \in [0.12, 0.14) \\ 3.71826 + v(1.00061 + (0.493797 + 0.193641v)v) & \text{if } v \in [0.14, 0.16) \\ \vdots & \\ 3.71532 + v(1.02469 + (0.424022 + 0.26667v)v) & \text{if } v \in [0.46, 0.48) \\ 3.71472 + v(1.02841 + (0.416264 + 0.272057v)v) & \text{if } v \in [0.48, 0.50) \\ 3.71403 + v(1.03253 + (0.40802 + 0.277553v)v) & \text{if } v \in [0.50, 0.52) \\ 3.71324 + v(1.03708 + (0.399273 + 0.28316v)v) & \text{if } v \in [0.52, 0.54) \\ \vdots & \\ 3.67497 + v(1.19609 + (0.175502 + 0.389949v)v) & \text{if } v \in [0.84, 0.86) \\ 3.66996 + v(1.21357 + (0.155178 + 0.397826v)v) & \text{if } v \in [0.86, 0.88) \\ 3.66448 + v(1.23224 + (0.133961 + 0.405863v)v) & \text{if } v \in [0.88, 0.90) \\ 3.65851 + v(1.25217 + (0.111823 + 0.414062v)v) & \text{if } v \in [0.90, 0.92) \\ 3.65199 + v(1.27341 + (0.0887367 + 0.422427v)v) & \text{if } v \in [0.92, 0.94) \\ 3.6449 + v(1.29603 + (0.0646717 + 0.430961v)v) & \text{if } v \in [0.94, 0.96) \\ 3.6372 + v(1.3201 + (0.0395983 + 0.439667v)v) & \text{if } v \in [0.96, 0.98) \\ 3.62884 + v(1.34569 + (0.0134853 + 0.448549v)v) & \text{if } v \in [0.98, 1.00). \end{cases} \tag{45}$$

Table 15 Absolute Error at Different Spatial Grid Points When $t = 1$, $N = 100$, $\gamma = 0.3$ and $\Delta t = 0.01$ of Example 13.

v	Exact Solution	Approximate Solution	Absolute Error
0.1	2.6402339273308577	2.6402334591110255	4.68220×10^{-7}
0.2	2.7564657674153796	2.7564649092297193	8.58186×10^{-7}
0.3	2.8849218168312127	2.8849206530242460	1.16381×10^{-6}
0.4	3.0268877068964803	3.0268863296990793	1.37720×10^{-6}
0.5	3.1837842799553380	3.1837827914642450	1.48849×10^{-6}
0.6	3.3571818096457187	3.3571803240160594	1.48563×10^{-6}
0.7	3.5488157167256860	3.5488143626082764	1.35412×10^{-6}
0.8	3.7606039377476774	3.7606028610063826	1.07674×10^{-6}
0.9	3.9946661204121600	3.9946654871604155	6.33252×10^{-7}

Table 16 Error Norms When $\Delta t = 0.01$, $\gamma = 0.6$ and $N = 64$ for Example 13.

t	$e_\infty(h, \Delta t)$	$e_2(h, \Delta t)$
0.2	2.17531×10^{-6}	1.57694×10^{-6}
0.4	7.57620×10^{-7}	4.94835×10^{-7}
0.6	1.34273×10^{-6}	9.40009×10^{-7}
0.8	3.69810×10^{-6}	2.69277×10^{-6}
1.0	6.83135×10^{-6}	4.99905×10^{-6}

Table 17 Convergence Order for Example 13 in Temporal Direction Where $h = \frac{1}{100}$ and $\gamma = 0.5$.

Δt	FDM ⁴⁷		Proposed Method		
	$e_\infty(h, \Delta t)$	Order	$e_\infty(h, \Delta t)$	$e_2(h, \Delta t)$	Order
$\frac{1}{10}$	2.21374×10^{-4}	...	2.13309×10^{-4}	1.56383×10^{-4}	...
$\frac{1}{20}$	5.02385×10^{-5}	2.13962	5.21507×10^{-5}	3.82320×10^{-5}	2.03219
$\frac{1}{40}$	1.24729×10^{-5}	2.07000	1.18261×10^{-5}	8.66785×10^{-6}	2.14071
$\frac{1}{80}$	2.92960×10^{-6}	2.09002	1.75263×10^{-6}	1.27695×10^{-6}	2.75438

Table 18 Convergence Order for Example 13 in Spatial Direction Where $\Delta t = \frac{1}{100}$ and $\gamma = 0.5$.

h	FDM ⁴⁷		Proposed method		
	$e_\infty(h, \Delta t)$	Order	$e_\infty(h, \Delta t)$	$e_2(h, \Delta t)$	Order
$\frac{1}{4}$	9.88048×10^{-4}	...	1.01779×10^{-3}	7.47137×10^{-4}	...
$\frac{1}{8}$	2.48469×10^{-4}	1.99152	2.51478×10^{-4}	1.85053×10^{-4}	2.01694
$\frac{1}{16}$	2.66470×10^{-5}	1.98000	6.16166×10^{-5}	4.50280×10^{-5}	2.02904
$\frac{1}{32}$	1.57142×10^{-5}	1.99000	1.38100×10^{-5}	1.00735×10^{-5}	2.15761

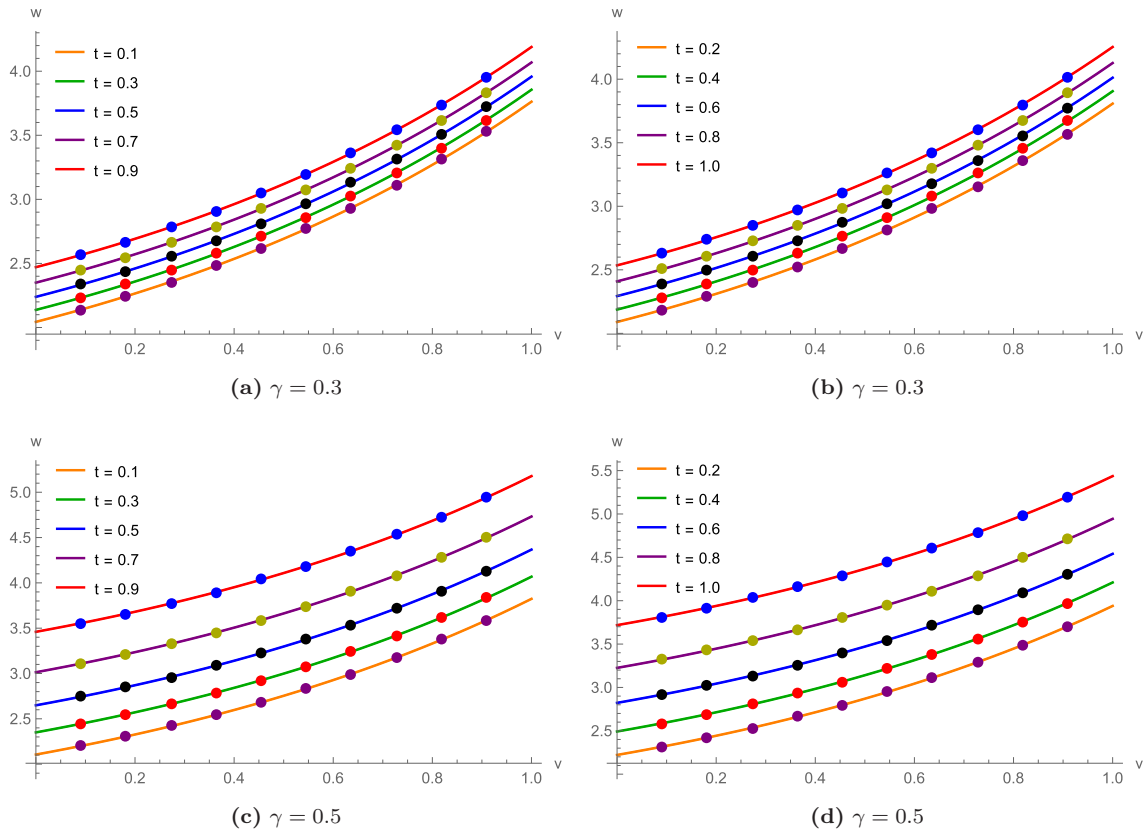


Fig. 9 Exact and approximate solutions at various time stages with $\Delta t = 0.01$ and $N = 32$ for Example 13.

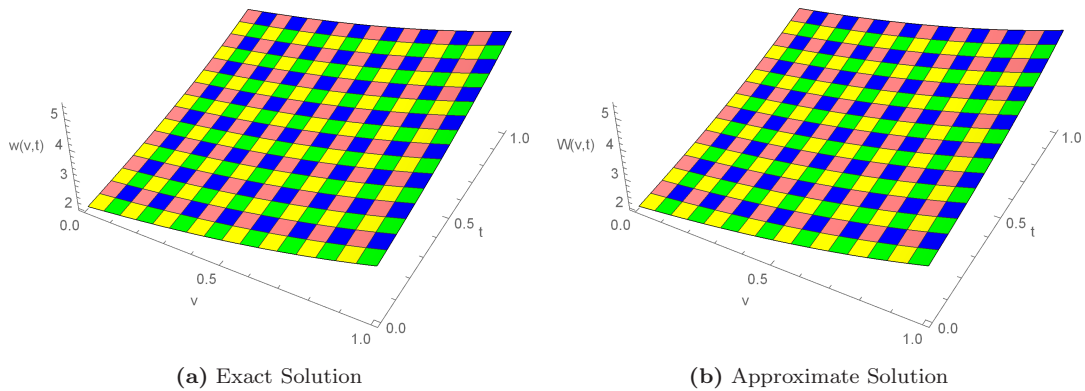


Fig. 10 3D exact and approximate solutions when $N = 90$, $\gamma = 0.5$, $t = 1$, $\Delta t = 0.01$ and $v \in [0, 1]$ for Example 13.

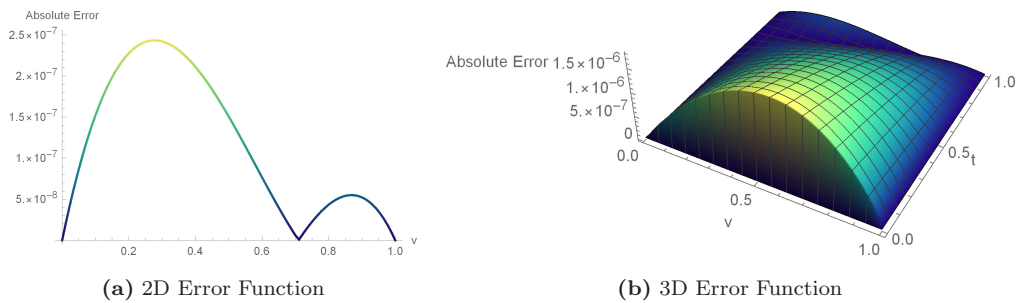


Fig. 11 2D and 3D error profiles when $N = 90$, $t = 1$, $\gamma = 0.5$, $\Delta t = 0.01$ and $v \in [0, 1]$ for Example 13.

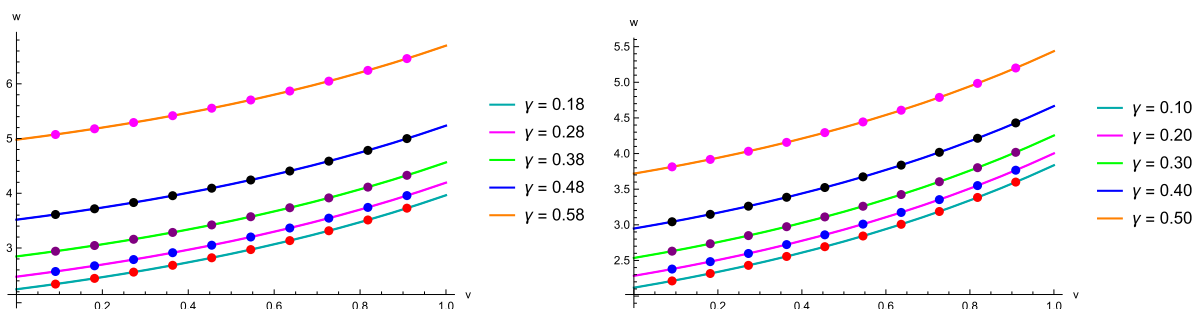


Fig. 12 Exact and approximate solutions at different γ values when $\Delta t = 0.01$, $N = 16$ and $t = 1$ of Example 13.

7. CONCLUSION

In this paper, an efficient numerical scheme based on CBS functions has been presented for solving TFDE involving Caputo–Fabrizio fractional time derivative. The non-integer order time derivative is approximated through the standard finite difference formulation while CBS functions are used to interpolate the solution curve in spatial direction. The current numerical algorithm is shown to be unconditionally stable and has second-order spatial and temporal convergence. The approximate results reveal that the presented method is reliable enough and efficient to be employed for solving mathematical models involving time fractional derivatives in Caputo–Fabrizio sense.

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